



# Time-integrated optical and extreme ultraviolet emission studies of the laser produced neon and nitrogen plasmas

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## Outline

- Motivation and aims
- Experimental setup and theoretical simulations
- Results of investigations
- Summary and outlook

















### **Motivation**

- Plasmas can be produced by photoionization of atomic and molecular gases with the use of intense EUV nanosecond pulses
- Spectral investigations of plasmas provide information about the kinetics and plasma parameters
- Research on photoionized plasmas is applicable in astrophysics and technology

### Aims

• To study atomic processes in photoionized plasmas created by LPP EUV/SXR sources.



## EXTATIC Schematic view and Experimental source First source





### Schematic view and Experimental source Second source





#### Laser plasma EUV source driven by a 0.8 J/4 ns/10 Hz Nd:YAG laser





EXTATIC















### **Gas Puff-target**









#### Measured and simulated emission spectra of Ne photoionized plasma



















#### Emission spectra emitted from Ne I in the visible range







#### Part of emission spectra observed in the ultraviolet region of Ne II







The slope of the best fit straight line indicated temperature of 1.8 eV in Ne II and almost 0.7 eV in Ne I









#### Spectra of gases ionized with EUV radiation pulses from a laser-plasma source: Neon



✓ Mostly, the spectral lines corresponding to

 $2s^2 2p^w nl - 2s^2 2p^{w-1} nl$  transitions

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## **Results Continue...**





















#### Examination of excitation and radiative rates in plasma: To clarify the LTE condition

The excitation rate coefficient  $\langle \sigma_{ik} v_e \rangle$  can be calculated as\*

$$\langle \sigma_{ik} v_e \rangle = 1.6 \times 10^{-5} \left( \frac{f_{ik} \langle \overline{g} \rangle}{\Delta E (kT_e)^{1/2}} \right) \exp \left( -\frac{\Delta E}{k_B T_e} \right)$$

The electron impact excitation rate is given by\*\*

$$S = n_e \langle \sigma_{ik} v_e \rangle$$

- $\langle \sigma_{ik} v_e \rangle$  represents coefficient rate of excitation
- $f_{ik}$  is represents the absorption oscillator strength
- $\langle \bar{g} \rangle$  is effective grant factor averaged over all Maxwellian velocity distribution function.
- $\Delta E$  (in eV) is the energy difference between the states and
- $k_{\rm B}$  and  $T_{\rm e}$  are respectively the Boltzmann constant and plasma electron temperature

\*H.J. Kunze, Space Sci. Rev. 13, 565 (1972)
\*\* J. D. Hey, J. Quant. Spectrosc. Radiat. Transf. 16, 69 (1976)

















#### **Rates of excitation and radiative decay**

Single charged ions	
Ne II 333.484 nm,	$2s^2 2p^4 3s \ 4P_{5/2} - 2s^2 2p^4 3p \ 4D_{7/2}^0$
Ne II 356.585 nm,	$2s^2 2p^4 3p \ 4S^0_{3/2} - 2s^2 2p^4 3d \ 4P_{3/2}$
Ne II 371.313 nm,	$2s^2 2p^4 3s 2P_{3/2} - 2s^2 2p^4 3p 2D_{5/2}^0$
Neutral atoms	
Ne I 584.258 nm,	$2s^2 2p^5 3s 2[1/2]_1^0 - 2s^2 2p^5 3p 2[1/2]_0$
Ne I 588.189 nm,	$2s^2 2p^4 3s 2[3/2]_2^0 - 2s^2 2p^4 3p 2[1/2]_1$

Ne I 594.483 nm,  $2s^2 2p^4 3s 2[3/2]_2^0 - 2s^2 2p^4 3p 2[3/2]_1$ 

















#### **Results of investigation for excitation and radiative decay rates**

Excitation rate coefficient			Excitation rate		10 times Radiative decay	
ion	line	Excita. Rate	n <sub>e</sub>	n <sub>e</sub> *Excita.	A <sub>ki</sub>	10*A <sub>ki</sub>
Ne II	333.484	2.726E-07	1.301E+15	3.5E+08	1.80E+08	1.80E+09
Ne II	356.585	2.972E-07	1.434E+15	4.3E+08	6.20E+07	6.20E+08
Ne II	371.313	3.688E-07	3.713E+15	1.4E+09	1.30E+08	1.30E+09
Ne I	585.258	2.650E-08	2.856E+16	7.6E+08	6.82E+07	6.82E+08
Ne I	588.189	8.273E-09	5.412E+16	4.5E+08	1.15E+07	1.15E+08
Ne I	594.483	1.440E-08	6.495E+16	9.4E+08	1.13E+07	1.13E+08

- ✓ The plasma electron density  $n_{\rm e}$  each of these lines has been evaluated from a Voigt's FWHM fit of the experimental profile
- ✓ Partial local thermodynamics equilibrium found to be realized













#### Measured and simulated UV/VIS emission spectra of N<sub>2</sub> photoionized plasma



LIFBASE: \*J. Luque and D.R. Crosley. Lifbase: Database and spectral simulation program (version 2.1.1). SRI International Report MP 99-009, 1999. Specair: E. Pawelec, Eur. Phys. J. Special Topics 144, 227-231 (2007)



















- ✓ Direct fitting of vibration bands give accurate data on the electron vibrational temperature
- ✓ And also result in a good agreement with the measured vibrational levels















DCI



Determination of vibrational  $T_v$  temperature of a plasma created in a molecular nitrogen



Vibrational populations of non-thermal mode vibrational levels of  $N_2^+$ , obtained at the end of the simulations for two different temperatures  $(T_v, T_r)$ .

Vibrational temperature determination from the Boltzmann plot of the  $\Delta v = 0, \pm 1, -2$ , and -3 sequences in second positive band transitions  $N_2(C^3\Pi_u - B^3\Pi_g)$ .

\*C.O. Laux, Radiation and Nonequilibrium Collisional-Radiative Models, von Karman Institute Lecture Series, July 2002.







- ✓ LLP EUV sources were used to study time-integrated emission lines from photoionized plasmas in atomic and molecular gases.
- ✓ The observed and identified spectral lines were mostly originated from radiative transitions in single and double charged ions.
- ✓ Initial calculations of electron impact excitation and radiative decay rates were presented from neutral and singly charged ions.
- ✓ We measured the band spectra (first negative and second positive systems) of the nitrogen molecule and we examined the vibrational temperature of the  $N_2(C^3\Pi_u B^3\Pi_g)$  band transitions.
- Collision frequencies between the plasma particles and other plasma parameters will be evaluated for hydrodynamics calculations of the LPP EUV source in the near future.





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# Thank you for your attention



















# **Approximate estimation of n**<sub>e</sub>

## Main free path of photoelectrons

 $l_{e}$ 

$$= \frac{1}{n\sigma} \qquad \qquad \sigma = \sigma_i + \sigma_{ex}$$
10<sup>-17</sup> cm<sup>2</sup>

Assume that thermalization is possible in case *l* <<d Where *d* is plasma size and assumed to be 1 mm

*n* ~ 2.25 x 10<sup>17</sup> cm<sup>-3</sup>







# The lower limit for electron density for which plasma, somehow to satisfy LTE\*

$$n_e(cm^{-3}) > 1.61018 \times 10^{12} \cdot T_e^{1/2} (\Delta E)^3$$

## $\Delta E$ is the higher energy different in eV of the levels whose populations are given by LTE conditions

\*R. W. P. McWhirter, Spectral Intensities, in Plasma Diagnostic Techniques, ed. R. H. Huddlestone and S. L. Leonard, Academic Press, New York, 1965, ch. 5, p. 206.















DC



Electron density has been determined from the line profiles of the isolated Krypton neutral lines neglecting the contribution of the ion impact broadening and Doppler broadening

$$\begin{split} &\Delta\lambda = 0.00884 \text{ nm, Lorentzian FWHM} \\ &\omega = 0.0886 \text{ nm, Stark broadening parameter} \\ &N_r \text{ (reference electron density).} \\ &\text{Neutral atoms: } N_r = 10^{16} \text{ cm}^{-3} \\ &\text{Singly ionized atoms: } N_r = 10^{17} \text{cm}^{-3} \end{split}$$

$$N_e = \frac{0.00884 \times 10^{16}}{2 \times 0.0886} \approx 5 \times 10^{14} \, cm^{-3}$$

N. Konjevic et al, J. of Phys. Chem. Ref. Data 31, 819 (2002).

\*J. Luque and D.R. Crosley. Lifbase: Database and spectral simulation program (version 2.1.1). SRI International Report MP 99-009, 1999.







## Abstract

Emission lines in different electromagnetic radiations from low temperature photoionized plasmas of atomic and molecular gases were investigated. The photoionized plasmas were induced using nanosecond radiation pulses of laser-produced plasma extreme ultraviolet (EUV) source. The source was based on a double stream gas puff target irradiated with a commercial Nd:YAG laser of 10 ns 10 Hz repetition rate with low and high pulse energy system. The EUV radiations were focused onto a gas stream, injected into a vacuum chamber synchronously with the EUV pulses. The radiation was collected and focused using different EUV collectors. A range of non-local equilibrium (NLTE) models and the radiative collisional codes have been used to reproduced theoretical spectra for further interpretation.











